Three-dimensional seismic velocity tomography of Montserrat from the SEA-CALIPSO offshore/onshore experiment

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[1] The SEA-CALIPSO experiment in December 2007 incorporated a sea-based airgun source, and seismic recorders both on Montserrat and on the adjacent sea floor. A high quality subset of the data was used for a first arrival P-wave velocity tomographic study. A total of more than 115,000 traveltime data from 4413 airgun shots, and 58 recording stations, were used in this high-resolution tomographic inversion. The experiment geometry limited the depth of well resolved structures to about 5 km. The most striking features of the tomography are three relatively high velocity zones below each of the main volcanic centers on Montserrat, and three low velocity zones flanking Centre Hills. We suggest that the high velocity zones represent the solid andesitic cores of the volcano complexes, characterized by wave speeds faster than adjacent volcaniclastic material. The low velocity zones may reflect porous volcaniclastic material and/or alteration by formerly active hydrothermal systems.


1. Introduction

[2] We present a P-wave velocity tomography study on Montserrat, W.I., based on the SEA-CALIPSO (Seismic Experiment with Airgun-source – Caribbean Andesitic Lava Island Precision Seismo-geodetic Observatory) offshore/onshore active source experiment. Montserrat is a 10 × 16 km volcanic island in the northern half of the Lesser Antilles arc. The Soufrière Hills volcano (SHV) dominates the southern two thirds of the island and has been active and dynamic since July 1995 [Young et al., 1998]. Various lines of evidence have been used to estimate the depth of over 5 km to a magmatic reservoir, including petrology [Barclay et al., 1998], deformation [Mattioli et al., 1998; Voight et al., 2006], and the approximate base of seismic activity [Aspinall et al., 1998]. We attempted to identify the regions of variable velocity under the island, especially under the volcanic centers, and to inquire whether magmatic storage areas could be recognized under the SHV.

2. Seismic Tomography Experiment

[3] The plan was to install a dense array of seismometers on or near the island, and to encircle the island with a ship towing an eight-component GI airgun with a total capacity of 2600 m³ shooting at 1 minute intervals (B. Voight et al., Active-source seismic experiment reveals structure under Soufriere Hills Volcano, Montserrat, WI, submitted to *Eos Transactons, American Geophysical Union*, 2010). A major focus concerned obtaining high resolution data under the active and hazardous volcano. Since SHV occupies the southeastern part of the island, most stations were located in safe areas in the north and northwest, and the ship’s radial tracks were mostly to the south and east (Figure 1). The short distance between the target zone and the furthest recorder limited the maximum depth of seismic ray penetration below SHV.

[4] Two deployment designs and types of seismic recorders were used on land in SEA-CALIPSO (Voight et al., submitted manuscript, 2010): 29 Reftek 130 recorders, with 3-component Mark Products L22 2 Hz sensors, and 204 Texans (Reftek 125), with single vertical component Mark Products L28 4.5 Hz sensors. In order not to bias the inversion with the closely spaced Texan reflection geometry, only one in ten Texans was included in the tomographic grid.

[5] The seismic network used in the tomographic inversion consisted of 58 stations, including 25 Reftek 130s, 19 Texans, 7 Ocean-Bottom Seismometers (OBSs) with 4.5 Hz sensors, and 7 permanent Montserrat Volcano Observatory (MVO) broadband stations in the exclusion zone. For the duration of the experiment, all instruments recorded continuously, at 250 or 100 samples per second depending on recorder type. The only exception was a subset of six Texans, part of a refraction line in the Belham Valley, that was shut down for several hours mid-experiment for a data quality test.

[6] The data recorded in the experiment was mostly of high quality with easily identifiable first arrivals (Figure 2). Weak signals were recorded from some of the longer ray paths. The first arrivals formed a smooth progression throughout the section; thus, even when a few shots did not show the first arrival, the arrival time could still be identified. In this study we analyzed and picked only the first
arrivals and did not use secondary phases. These phases are addressed in other works [Paulatto et al., 2010].

3. Data

[7] The 77 hours of shooting resulted in 4413 shots recorded by 58 stations for a total of 115,158 ray paths. Of all the shots, stations in the north had up to 91% of identifiable first arrivals, while stations around SHV had fewer than 35% identifiable first arrivals due to high attenuation in the younger volcanic deposits. Stations near the coast recorded few first arrivals due to high ambient noise from ocean waves. The data were filtered between 3–15 Hz to correspond with the frequency content of the airgun signal. First arrivals were picked manually and only clear signals were used; thus, uncertainty was assumed to be the same for all first arrivals.

4. Seismic Tomography Method

[8] The first-arrival time data were inverted for a 3D P-wave velocity model of Montserrat and the surrounding ocean using the tomography code from Shalev and Lees [1998]. This method uses a Cubic B-spline description of the 3D volume, and the LSQR algorithm to invert the data. This inversion method simultaneously minimized both data misfit and model roughness, which allowed the researcher to choose the desired level of smoothness. The inversion also allowed for station and shot corrections. A separate damping parameter was used for each type of inversion unknown: velocity model, station correction, and shot correction.

[9] Although the 3-D velocity structure extended to 1 km above sea level, ray tracing was computed from sources at sea level to the actual elevation of each station. The length of the water section of each path was assumed to be the depth to the sea floor below each shot. A constant velocity of 1.5 km/s was used to calculate travel time in the water.

[10] We began with a 1D velocity model, using Cubic B-spline interpolation to be consistent with the 3-D inversion. The 1D model for this study was derived from the data using the Levenberg–Marquardt non-linear minimization procedure [Press et al., 1992]. There were two options for the starting velocity model: a) A single starting model for the whole target area, or b) Two starting models, one for land and one for ocean. We tested both types of starting model using the same damping and smoothing parameters, and while the end results were similar, the final RMS of residuals was smaller with two starting models. Therefore, we derived two 1D velocity models, one each for land and for ocean (Figure 2). The boundary between land and ocean was defined as the bathymetric line at 200 m water depth. The derived 1D models differed from the ones used for synthetic ray-path modeling. In particular, the final land model was faster than the ocean model at all depths penetrated by first arrival rays; this resulted in a shallower maximum depth for the turning rays, and reduced the imaging depth.

[11] The total target cube for the 3D inversion (see Figure 1) was 50 × 45 × 8 km. Horizontal velocity grid spacing was 0.5 km in the land area, 1 km for the ocean near the land, and 5 km near the boundaries (see auxiliary material). Vertical grid spacing was 0.5 km to a depth of 5 km, and 1 km below 5 km. A smaller grid spacing of 0.25 km was tested for the center of the land area but showed no improvement.

[12] To check for resolution of the 3D inversion, we ran a checkerboard test based on the starting 1D velocity models. The cell size of the checkerboard was 1.5 × 1.5 × 1.5 km, with the specified variation ±12.5%. A consistent recovery of the pattern was observed to a depth of 4 km in the area of good ray coverage under the island. However, the amplitude of the recovered anomalies was only about two-thirds of the starting amplitude. At 5 km depth, some of the checkerboard anomalies retained their shape but most were blended and smeared. There was no reliable resolution below 5 km depth (see auxiliary material).

[13] Another resolution problem was the tradeoff between station correction and velocity in the top 1 km of the model. The process of allowing for station corrections in the inversion removed most of the variability from the top of the model, and running the inversion without station corrections resulted in a substantially larger residual RMS. The final inversion allows for damped station correction. Nevertheless,

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1Auxiliary materials are available in the HTML. doi:10.1029/2010GL042498.
it is possible that some velocity anomalies in the top 1 km are not imaged because of this compromise.

5. Results

The tomographic inversion converged after 5 iterations. Several sets of damping and smoothing parameters were tested, and the major velocity anomalies were stable regardless of these parameters. The 3D inversion reduced the RMS of the residuals from 167 ms to 80 ms for a variance reduction of 77%. Results of the tomographic inversion are shown in Figures 3 and 4. The maximum lateral variation is about 2 km s$^{-1}$, which is similar to the results obtained from Deception Island [Zandomeneghi et al., 2009].

The most notable features in the Vp structure are high velocity anomalies below all three volcanic centres of Montserrat at about 2 to 3 km depth. The most prominent of these is the anomaly below Centre Hills, with a similar but less intense anomaly under SHV (Figure 3a). This inversion result was not unexpected, because the average station residual contours in Figure 1 also show positive (fast) residuals for stations near the volcanic centers. The top boundaries of these high velocity anomalies are not clear, due to the tradeoff between station correction and velocity at shallow depth. When the inversion ran without station corrections, the largest fast anomaly under Centre Hills reached the surface, but with higher RMS misfit.

Other large and consistent anomalies are the low velocity regions on the flanks of the volcanic centres. There are three such anomalies to the northeast, northwest, and southwest of Centre Hills. These anomalies are stable regardless of inversion parameters. The east-west cross section (Figure 3d) shows both a high velocity body under SHV and a low velocity anomaly west of SHV. The appearance that the two anomalies are elongated down and away from the center of the island may be an artefact of the rays coming from the perimeter to the center. However, the geometry of the main, high amplitude anomalies of Figure 4 is stable.

6. Discussion

The acquisition geometry of the SEA-CALIPSO tomographic experiment was laid out to target the possible active concentration of magma at >5 km depth under the SHV. The actual seismic velocities beneath and surrounding Montserrat turned out to be faster than expected, thus
turning back most of the refracting seismic energy at depths $<5$ km. The result was that first-arrival P-wave tomography produced a reliable image of the velocity structure between $\sim1$ and 5 km in depth and extending approximately to the shelf break.

[18] Within this region are six prominent velocity anomalies enclosing perturbations either 6% above or below the average velocities at each depth (Figure 4). Following Paulatto et al. [2010], we suggest that the fast anomalies beneath the three constructional volcanic centers may correspond to solid andesitic structural elements in the volcanic cores. The cores would consist of dense, crystallized rock comprising dome cores, sills, dikes, or irregular-shaped intrusions, and adjacent altered zones with silica precipitation, that are seismically faster than the surrounding material, the latter including either lavas from submarine volcano building, and volcaniclastic deposits (talus, block-and-ash flows, lahars etc.). Crystalline cores are consistent with the work of Harford and Sparks [2001], who suggest that recurring intrusions solidify at depths up to $\sim3$ km under SHV. This is supported by other evidence that suggests that dikes may rise to 1.5–2 km under SHV, from shallow storage zones [Mattioli et al., 1998; Hautmann et al., 2009; Voight et al., 2010]. It is likely that intrusions have some lateral extent [e.g., Voight et al., 2006], and that considerable volumes of unerupted magma remain in storage zones during the current eruptive activity. The high velocities observed are consistent with nodules found Montserrat-wide as inclusions in eruption products [Kiddle et al., 2010].

[19] The locations of the low-velocity anomalies northeast of Centre Hills and west of SHV suggest a relationship with the volcanic centers and the features may represent syn-volcanic apron deposits. A potential weakness with this hypothesis lies in the fact that these low-velocity features extend to 3–4 km depth. While the weight of the volcano could have been responsible for some degree of crustal flexure and burial of surface material, it cannot easily account for such a depth. There is evidence for buried volcaniclastic
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References

Figure 4. Three-dimensional isosurfaces of velocity anomalies. The blue surfaces define anomalies that are >6% faster than average. The red surfaces represent anomalies that are >6% slower than average. (a) A map view. (b) A view from the east southeast. (C) A view from the south southwest.


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